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12.510 Introduction to Seismology Spring 2008

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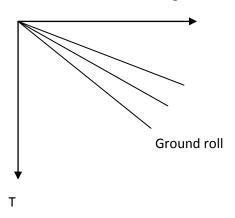
Continents: Quick review. Surface waves

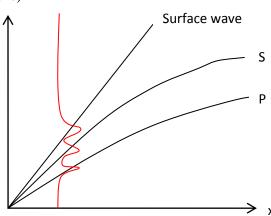
Ground roll-acoustic $\ddot{p} = V. \nabla. \left(\frac{1}{p} \nabla p\right)$. where p is the pressure

Love waves – SH $\ddot{u}_y = \mu . \nabla . \left(\frac{1}{p} \nabla u_y\right)$

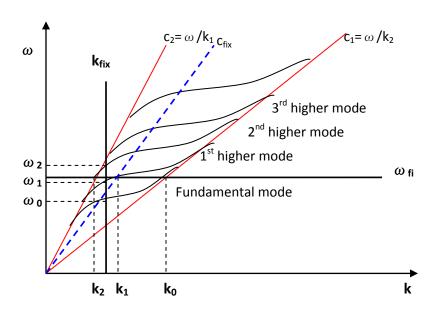
Rayleigh waves P-SV

Quick review (refer to April, 4, 2008 for details)

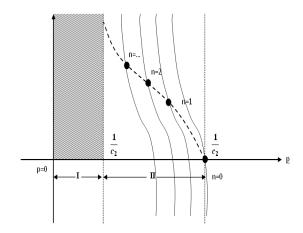




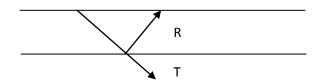
ω – k domain



Ground roll dispersion relationship

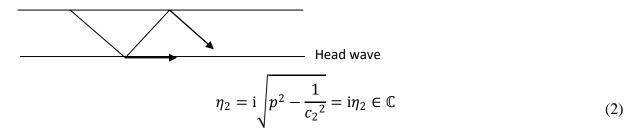


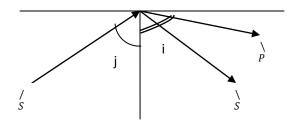
Pre-cretical $p < \frac{1}{c_2}$



$$\eta_2 = \sqrt{\frac{1}{{c_2}^2} - p^2} \in \mathbb{R} \tag{1}$$

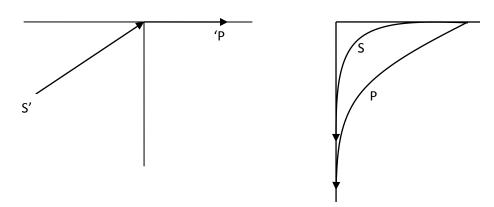
Post-critical $p > \frac{1}{c_2}$





$$\frac{\sin(j)}{\beta} = \frac{\sin(i)}{\alpha} = p \tag{3}$$

if $\beta < \alpha$, there can be critical reflection and horizontal propagating p-wave. If $j > j_c$ then there will be evanescence in the p-wave ($p > 1/\alpha$).



Both p-wave and s-wave horizontal propagation if $p > \frac{1}{\beta} > \frac{1}{\alpha} = \frac{1}{c}$. If a wave comes in with a

1/c that is larger than local $1/\alpha$ and $1/\beta$, the above will occur. This will also happen if the source emits a horizontal energy (rare).

$$P: \emptyset = Aexp(-\omega \eta_{\alpha} z) exp(i\omega(px-t)), \eta_{\alpha} = \sqrt{\frac{1}{\alpha^{2}} - p^{2}} \underset{p>\frac{1}{\alpha}}{=} i\sqrt{p^{2} - \frac{1}{\alpha^{2}}} = i\widehat{\eta_{\alpha}} = i\sqrt{\left(\frac{1}{c}\right)^{2} - \left(\frac{1}{\alpha}\right)^{2}}; \text{ where } c = c_{R} = \frac{1}{p}(c_{R}: \text{ phase velocity for Rayleigh waves})$$

$$(4)$$

$$S: \psi = \beta exp(-\omega \eta_{\beta} z) exp(i\omega(px - t)), \eta_{\beta} = \sqrt{\frac{1}{\beta^{2}} - p^{2}} \underset{p > \frac{1}{\beta}}{=} i \sqrt{p^{2} - \frac{1}{\beta^{2}}} = i \widehat{\eta_{\beta}}$$

$$= i \sqrt{\left(\frac{1}{c}\right)^{2} - \left(\frac{1}{\beta}\right)^{2}}$$
(5)

We follow the same "Recipe" we used before:

- 1. Potentials
- 2. Boundary Conditions (Kinematic and dynamic)
- 3. Zoeppritz equatios.

Boundary conditions

$$u(x,t) = \nabla \emptyset + \nabla \times \psi \tag{6}$$

In this case

After some work we get:

$$A[(\lambda + 2\mu)\eta_{\alpha}^{2} + \lambda p^{2}] + 2\mu p \eta_{\beta} = 0$$

$$A(2p\eta_{\alpha}) + B(p^{2} - \eta_{\beta}^{2}) = 0$$
(9)

Zoepprtiz

$$\begin{bmatrix} (\lambda + 2\mu)\eta_{\alpha}^{2} + \lambda p^{2} & 2\mu p\eta_{\beta} \\ 2p\eta_{\alpha} & p^{2} - \eta_{\beta}^{2} \end{bmatrix} \begin{bmatrix} A \\ B \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \end{bmatrix}$$
 (10)

Trivial solution is

$$A = B = 0$$

Non-trivial solution leads to:

$$[(\lambda + 2\mu)\eta_{\alpha}^{2} + \lambda p^{2}][p^{2} - \eta_{\beta}^{2}] - 2p\eta_{\alpha}(2\mu p\eta_{\beta}) = 0$$
(11)

This expression; Rayleigh wave denominator (Rayleigh, 1887), can be written looking at wave speeds, but usually done numerically assuming a Poisson's medium ($\lambda=\mu$) and $\alpha=\sqrt{3}\beta$. This scaling will help to simplify the above equation.

$$\eta_{\alpha} = \sqrt{\frac{1}{\alpha^2} - p^2} = \sqrt{\frac{1}{\alpha^2} - \left(\frac{1}{c}\right)^2}$$

$$\eta_{\beta} = \sqrt{\frac{1}{\beta^2} - p^2} = \sqrt{\frac{1}{\beta^2} - \left(\frac{1}{c}\right)^2}$$
(12)

$$\alpha = \sqrt{\frac{\lambda + 2\mu}{\rho}}$$

$$\Rightarrow \rho\alpha^2 = \lambda + 2\mu$$

$$\beta = \sqrt{\frac{\mu}{\rho}}$$

$$\Rightarrow \rho \beta^2 = \mu$$

$$\left[\alpha^2 \left(\frac{\eta_{\alpha}^2}{\rho^2} + 1\right) - 2\beta^2\right] \left(1 - \frac{\eta_{\beta}^2}{\rho^2}\right) - \left(\frac{4\beta^2 \eta_{\alpha} \eta_{\beta}}{\rho^2}\right) = 0$$

$$\Rightarrow \frac{\beta}{\alpha}$$

Assumptions

Poisson medium: $\lambda = \mu$

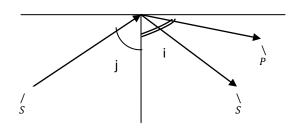
Poisson ratio: $V = \frac{\dot{\lambda}}{2(\lambda + \mu)} = 0.25$

$$\alpha = \sqrt{3}\beta$$

$$\left(\frac{c_R^2}{m{eta}^2}\right)^3 - 8\left(\frac{c_R^2}{m{eta}^2}\right)^2 + \frac{56}{3}\left(\frac{c_R^2}{m{eta}^2}\right) - \frac{32}{3} = 0$$

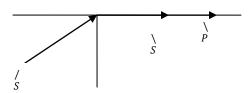
Three solutions

1.
$$\left(\frac{c}{\beta}\right)^2 = 4 \Rightarrow c = 2\beta, c > \beta, p = \frac{1}{c} < \frac{1}{p}, p < \frac{1}{\alpha}$$



2.
$$\left(\frac{c}{\beta}\right)^2 = 2 + \frac{2}{3}\sqrt{3}, \frac{1}{\alpha}$$

$$3.\left(\frac{c}{\beta}\right)^2 = (2 - \frac{2}{3}\sqrt{3}), \frac{1}{\alpha} \frac{1}{\beta} > \frac{1}{\alpha}, (\eta_{\alpha} = i\widehat{\eta_{\alpha}}, \eta_{\beta} = i\widehat{\eta_{\beta}})$$



Note: $c_R \sim 0.92\beta$ which makes it about 10% slower than the shear wave velocity. This explains why the Rayleigh wave is slower than the love wave. The Love wave, at low frequency $\rightarrow c_2$ and at high frequencies Love wave (at it's slowest) is c_1 ...Rayleigh wave is about 90% of the Love wave at the most.

exponential decay with depth

 u_x , u_z are $\frac{\pi}{2}$ out of phase

Updated by: Sami Alsaadan

Sources: April 13, 2008 by Patricia Gregg. April 8, 2008 lecture.

"An Introduction to Seismology, Earthquakes, And Earth Structure" by Stein & Wysession (2007).